

The Pan-STARRS search for Near Earth Asteroids - Present Status and Future Plans

Richard Wainscoat

Peter Vereš, Bryce Bolin, Marco Micheli, Larry Denneau, Robert Jedicke, Serge Chastel

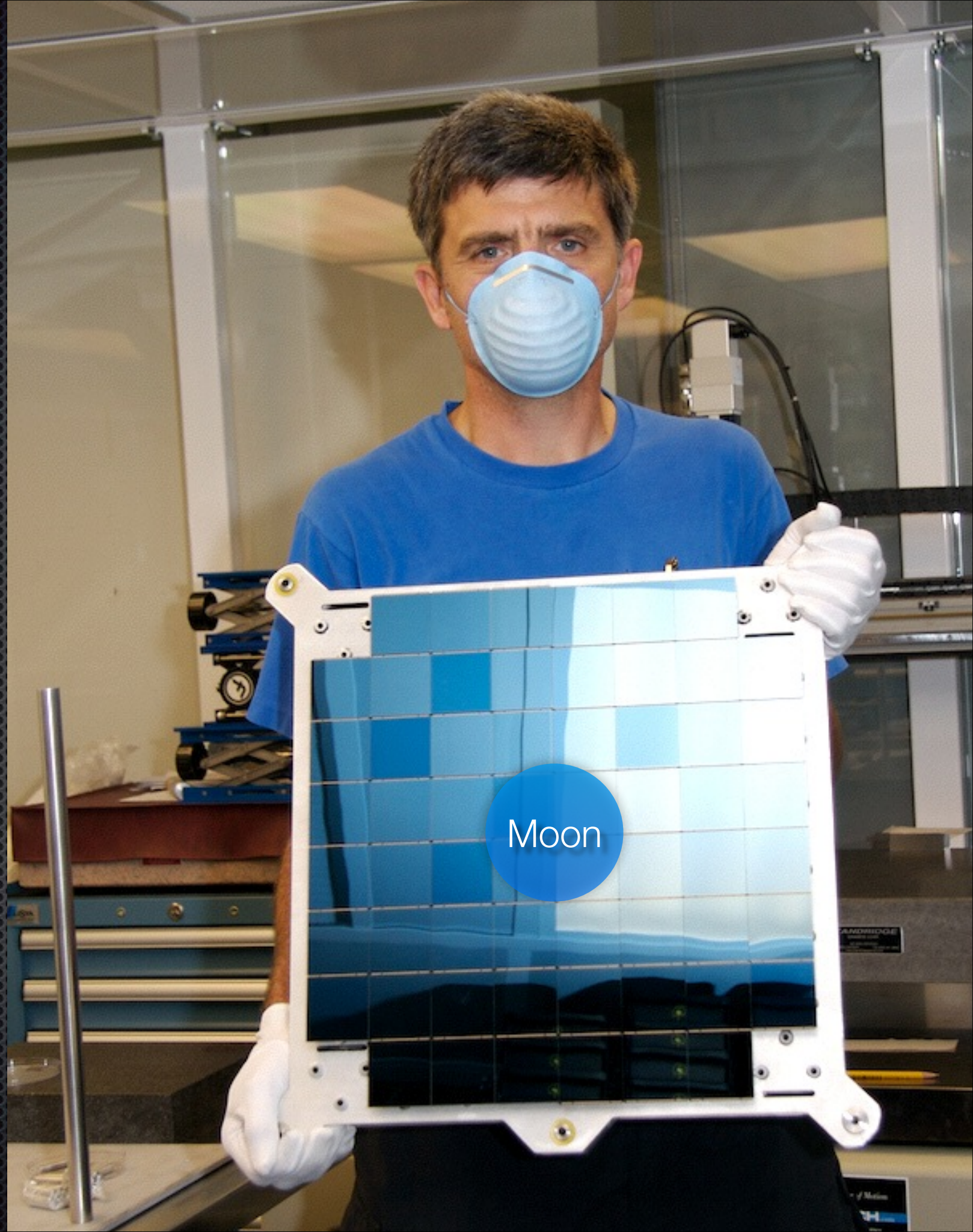
University of Hawaii, Institute for Astronomy

The Pan-STARRS telescopes

- ✦ Pan-STARRS1 is a 1.8-meter diameter telescope at Haleakala observatory in Maui (PS1)
- ✦ A second telescope (PS2) is being built adjacent to PS1
- ✦ PS1 has the largest digital camera in the world
 - ✦ 60 CCDs, each with 4800x4800 pixels, arranged in an 8x8 grid of 600x600 pixels

Gigapixel camera

- ✦ 1,382,400,000 pixels
- ✦ 7 square degree field-of-view
- ✦ Read time 12 sec
- ✦ Some CCDs are cosmetically poor
 - ✦ 70% fill factor



Pan-STARRS1 mission

- ✦ The observations that Pan-STARRS1 is obtaining are designed for many different scientific goals, including the solar system, brown dwarfs, Galactic structure, supernovae and other transients, and cosmology
- ✦ Nearly all observations that Pan-STARRS 1 obtains are executed in a manner that allows them to be searched for Near Earth Objects
- ✦ NASA is funding the NEO search with PS1
- ✦ Astrometry from PS1 is excellent, and in most cases is better than 0.15 arcsec

3pi survey

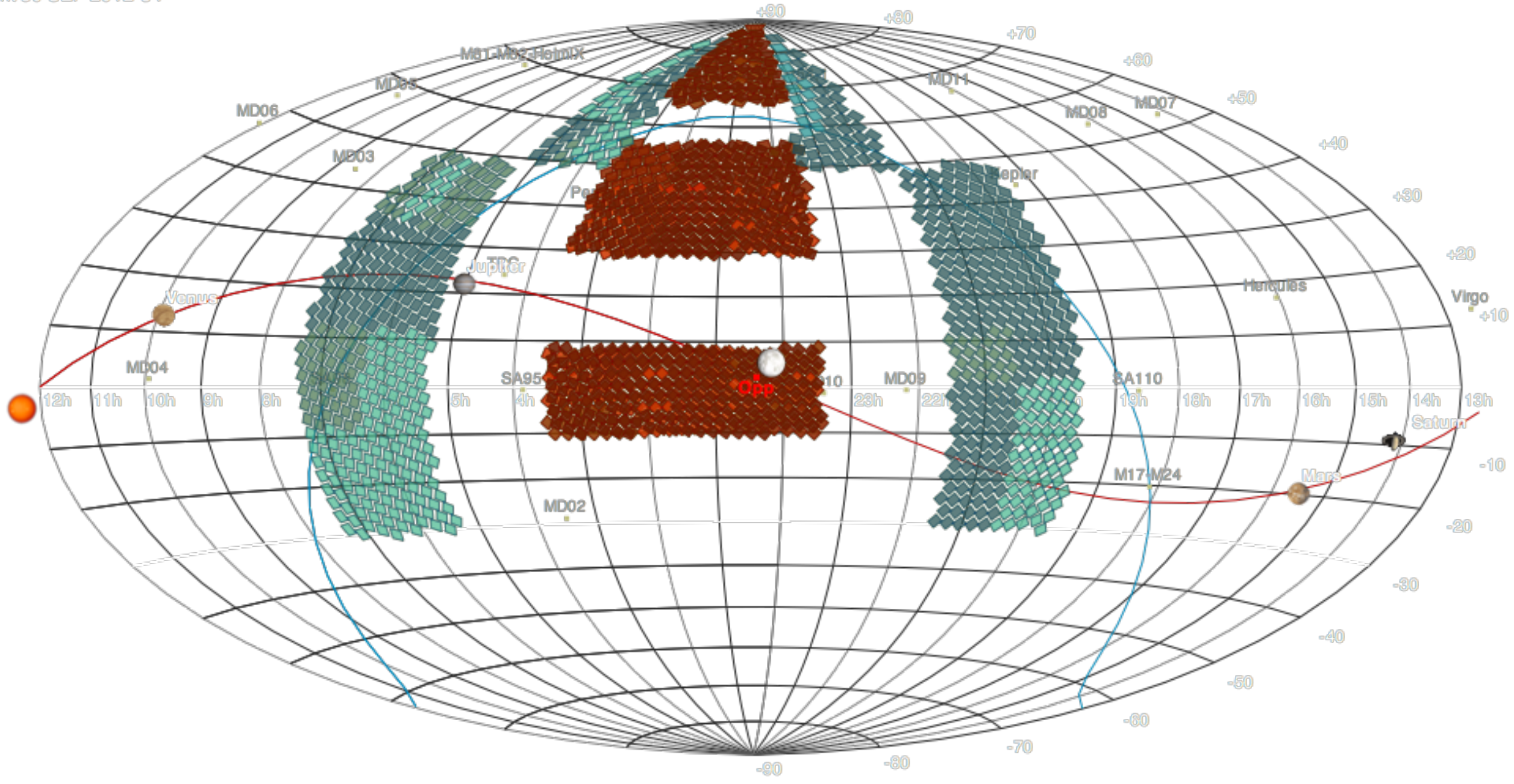
- ✦ Entire sky north of -30 surveyed every year
- ✦ 2,500 square degrees observed every month using 56% of the observing time
- ✦ Two pairs in each of g, r, and i (12 observations total)
- ✦ Each pair separated by approximately 20 minutes
- ✦ The entire area is observed at least once as a quad (four observations each spaced 20 minutes apart)
- ✦ 239 NEOs discovered to date including 24 PHAs

October 3pi fields

137 | 138 | 139 | 140 | 141 | 142 | 143 | 144 | 145 | 146 | 147 | 148 | 149 | 150 | 151 | 152 | 153 | 154 | 155 | 156 | 157 | 158 | 159

157

2012 UT to 28 OCT 2012 UT
beginning night 30 SEP 2012 UT



fields matching "3pi"
Mollweide | Spherical

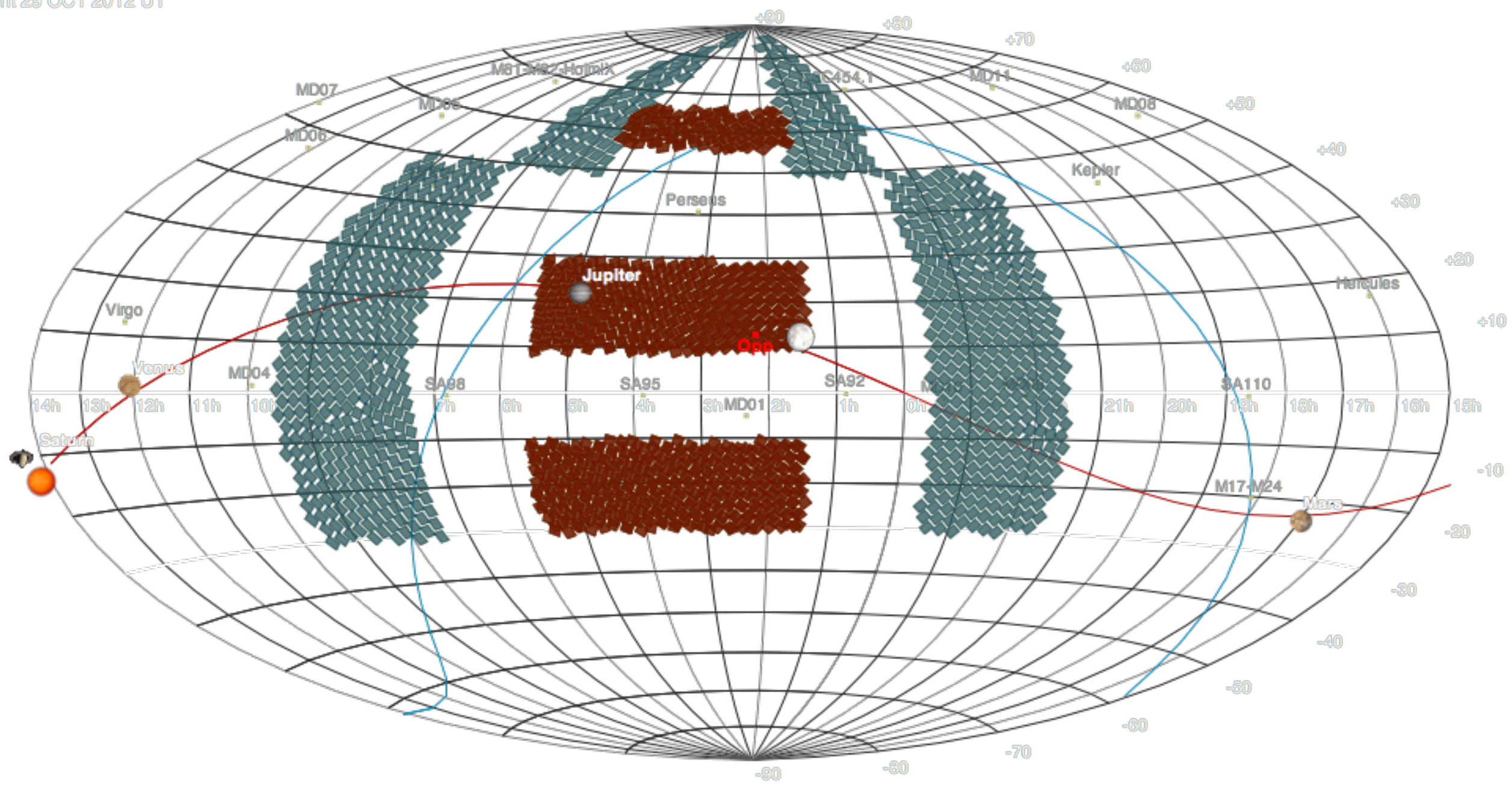
Click in plot to rotate

November 3pi fields

137 | 138 | 139 | 140 | 141 | 142 | 143 | 144 | 145 | 146 | 147 | 148 | 149 | 150 | 151 | 152 | 153 | 154 | 155 | 156 | 157 | 158 | 159

158

2012 UT to 27 NOV 2012 UT
beginning night 29 OCT 2012 UT



fields matching "3pi"

Wright-Aitoff | Spherical

Click in plot to rotate

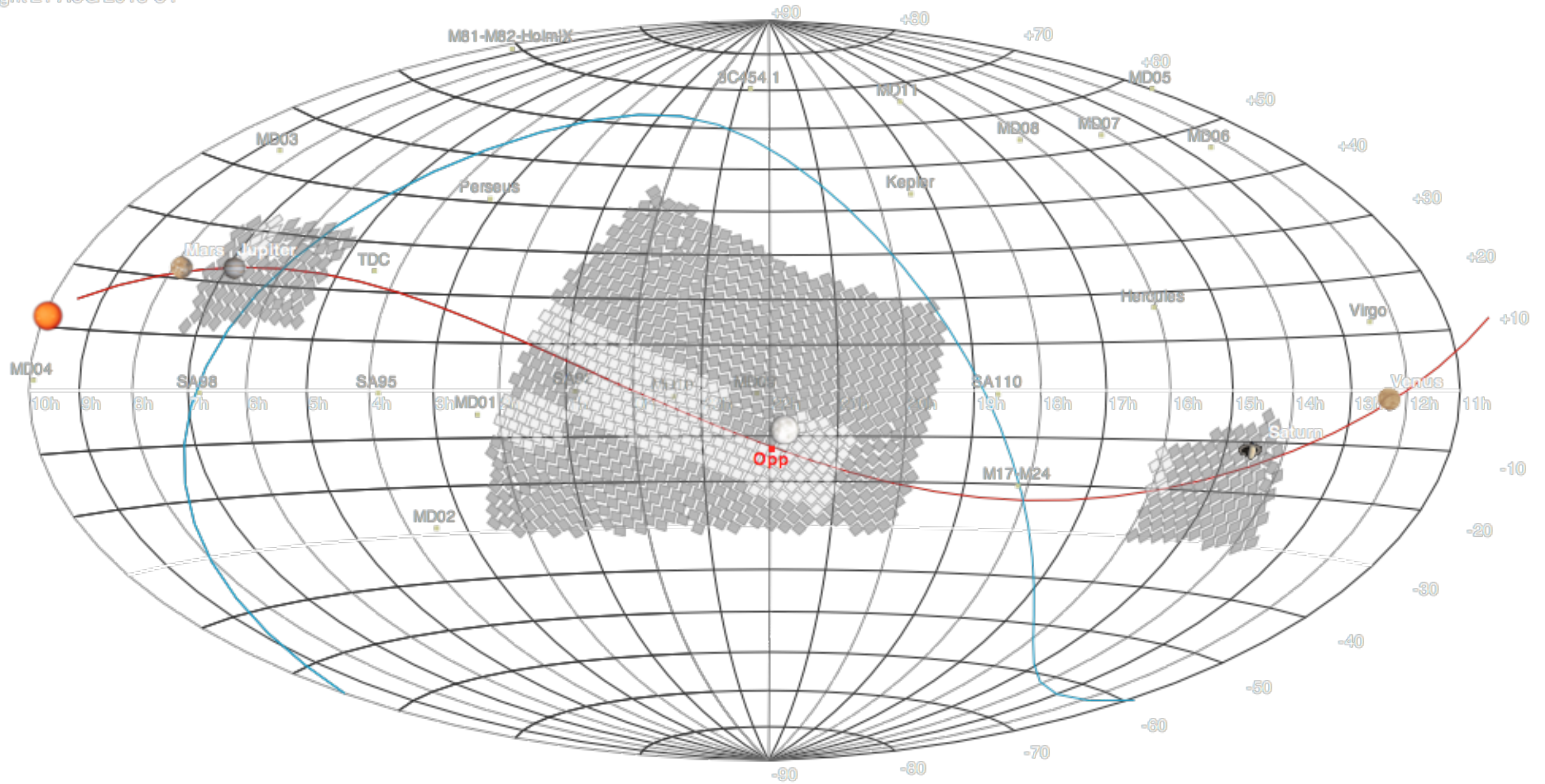
NEO Optimized survey

- ✦ We use a broad w-band filter (g+r+i) to increase sensitivity and have 11% of the observing time
- ✦ Four 45 second exposures separated by 20 minutes in the opposition direction or 7 minutes in the low solar elongation sweet spot directions, using wide filter
- ✦ Opposition search has yielded 536 NEO discoveries, including 41 PHAs
- ✦ Sweet spot search has yielded 28 NEO discoveries, including 11 PHAs

September 2013 solar system fields

168

2013 UT to 18 SEP 2013 UT
ending night 21 AUG 2013 UT



fields matching "ss"

Click in plot to

The need for followup observations

- ✦ Pan-STARRS NEO candidates are submitted to the Minor Planet Center
 - ✦ Pan-STARRS does not followup its own NEO candidates
 - ✦ We follow up some NEO candidates with CFHT
- ✦ PS1 is extremely dependent on other people to followup candidates to produce discoveries

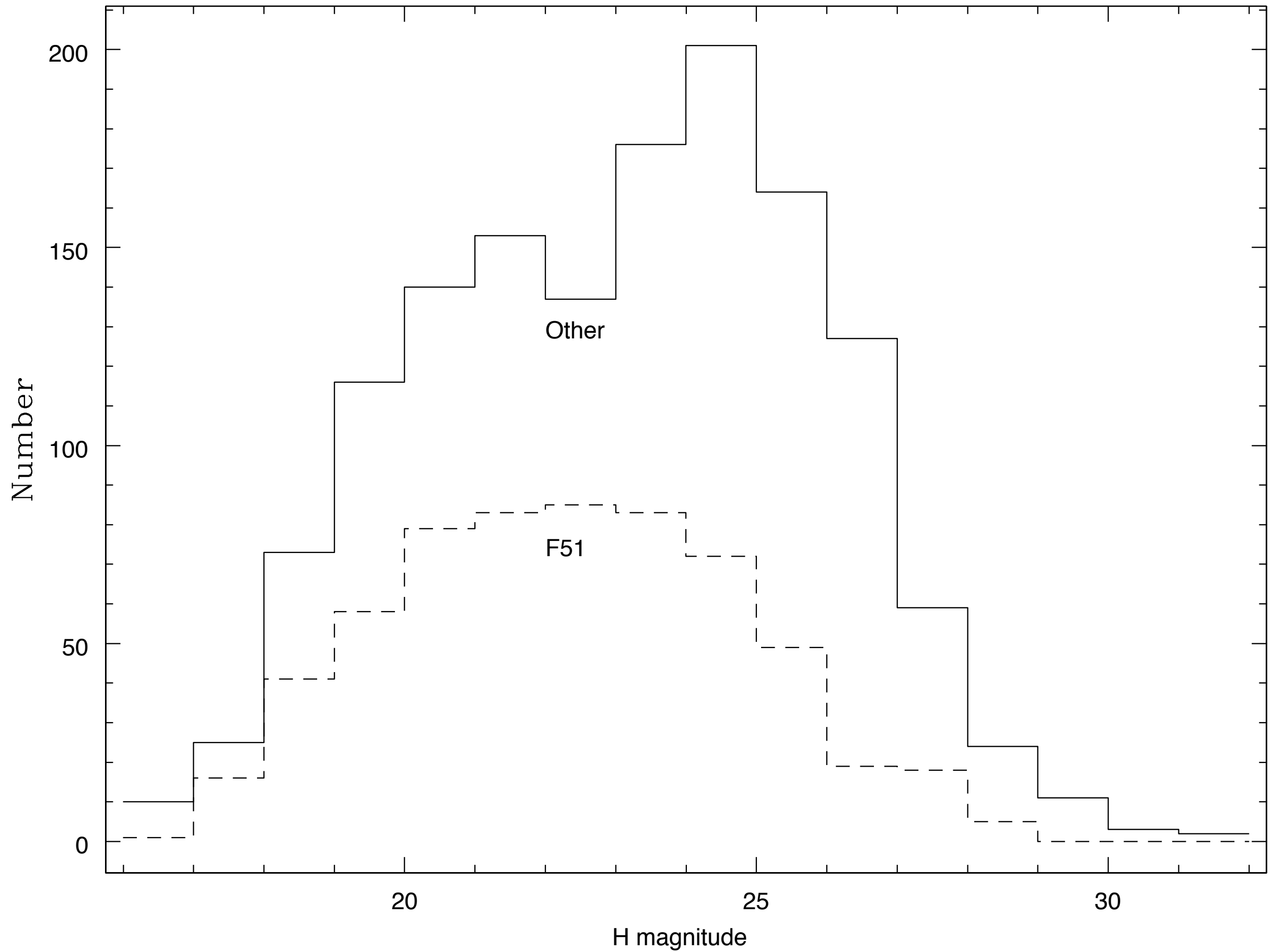
Submissions to MPC

- ✦ 2,279,348 asteroid tracklets reported to date
- ✦ 7,223,594 detections
- ✦ 563,836 distinct asteroids
- ✦ 40,357 asteroid discoveries
- ✦ 3,000–6,000 asteroids submitted per night

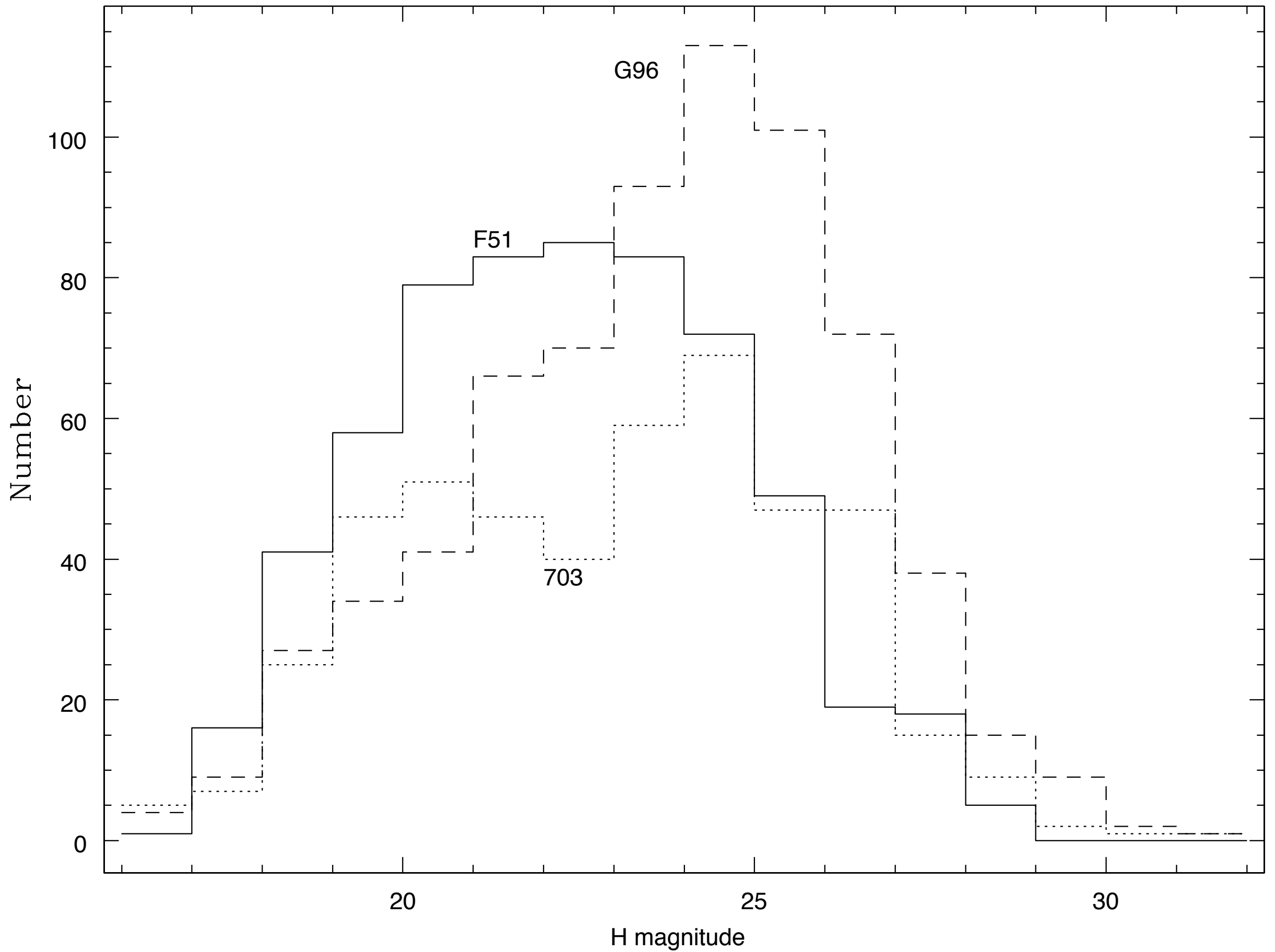
Discovery rate

- ✦ PS1 discovers over 50 NEOs per month when the weather is good
- ✦ The median H magnitude for PS1 NEO discoveries for 2012 and 2013 is 22.4
 - ✦ Other NEO discoveries (mostly from the Catalina Sky Survey) have median $H=23.4$
- ✦ Pan-STARRS is good at finding larger undiscovered NEOs that are distant and faint, but less efficient at finding smaller fast-moving nearby NEOs

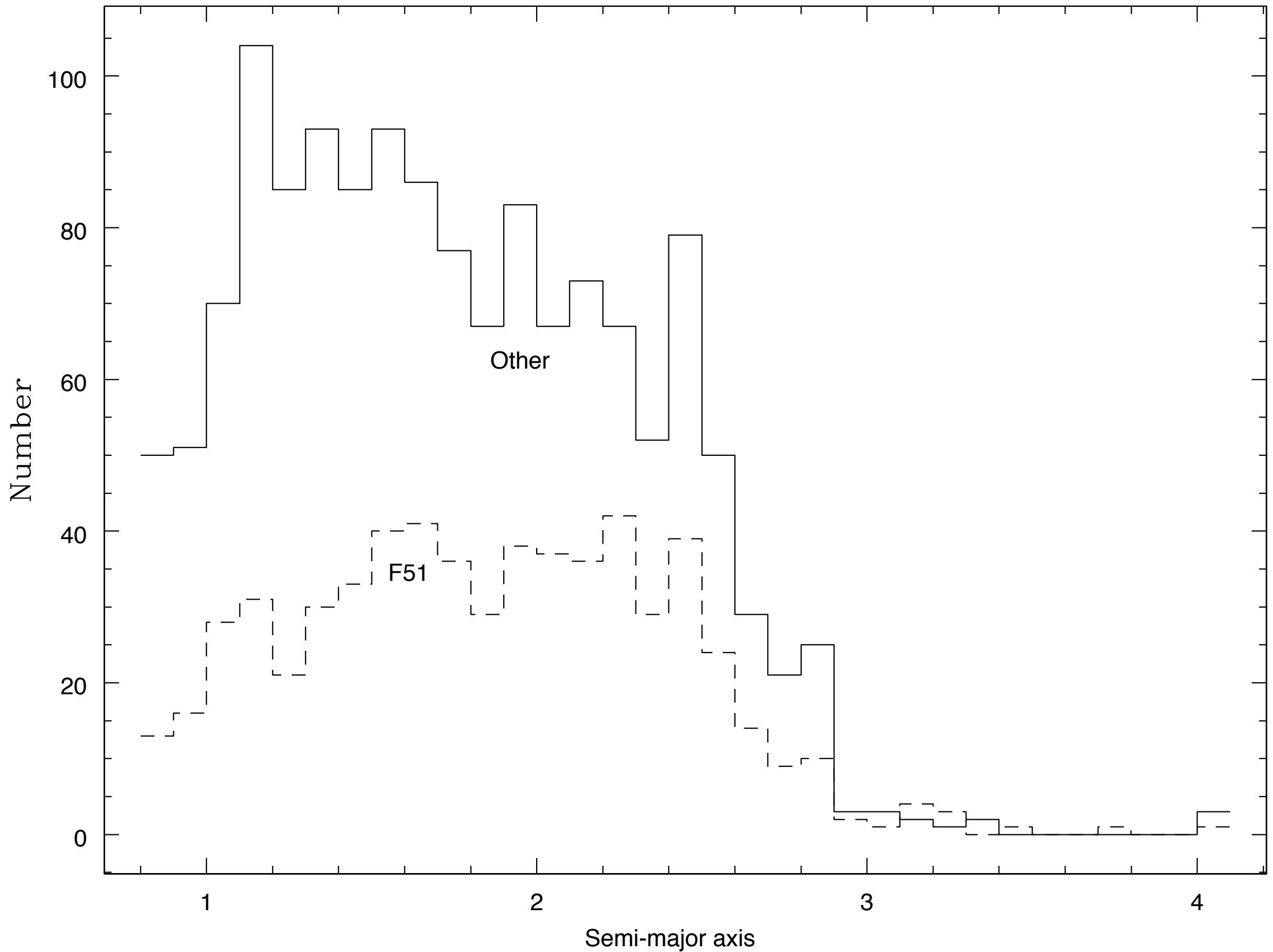
2012 and 2013 NEO discoveries



2012 and 2013 NEO discoveries



2012 and 2013 NEO discoveries



Comets

- ✦ Pan-STARRS has discovered many comets and is efficient at discovering low levels of activity
- ✦ Five Main Belt Comets have been discovered to date

Comet P2013/R3



✦ September

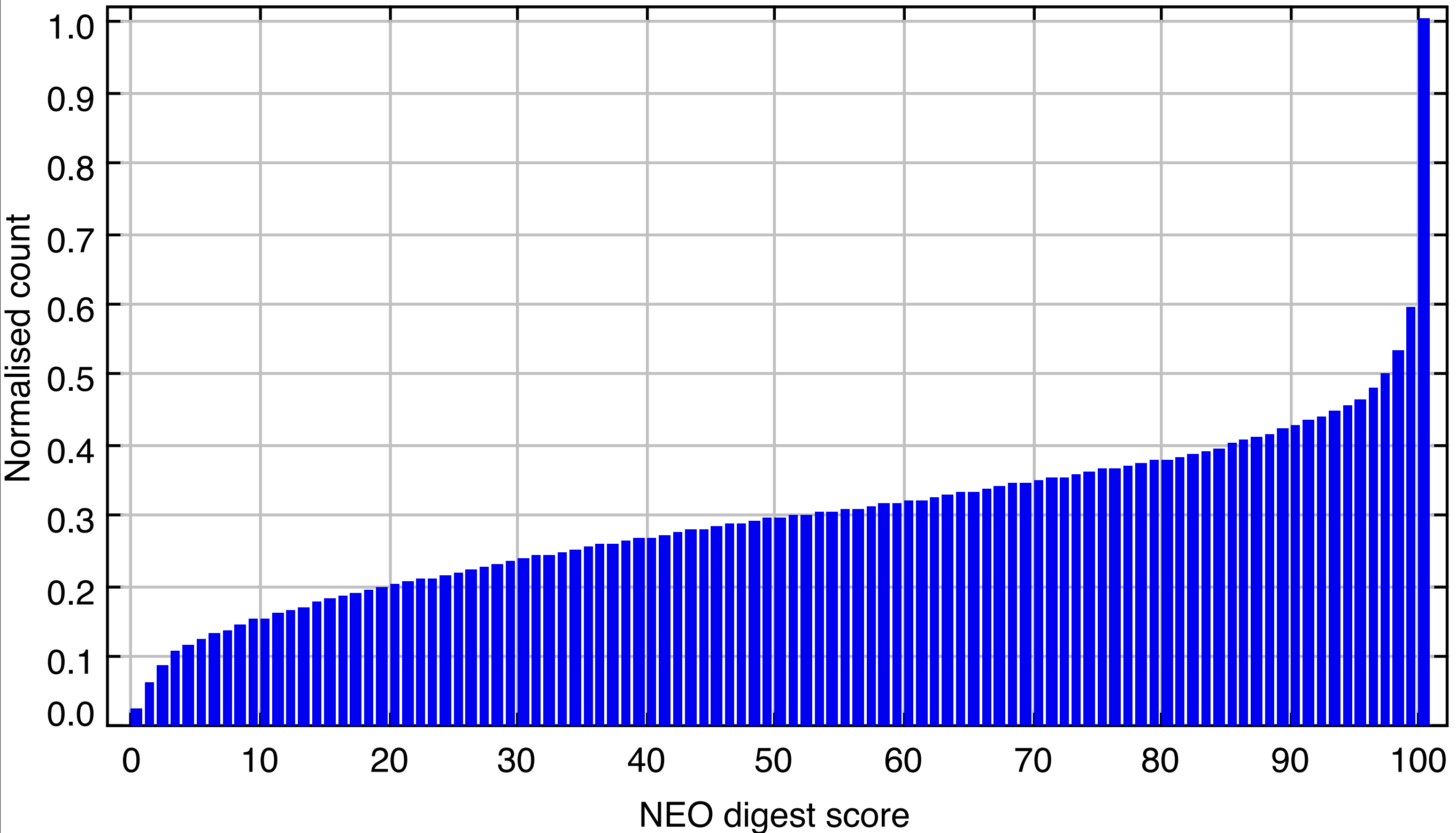
✦ October

Future plans

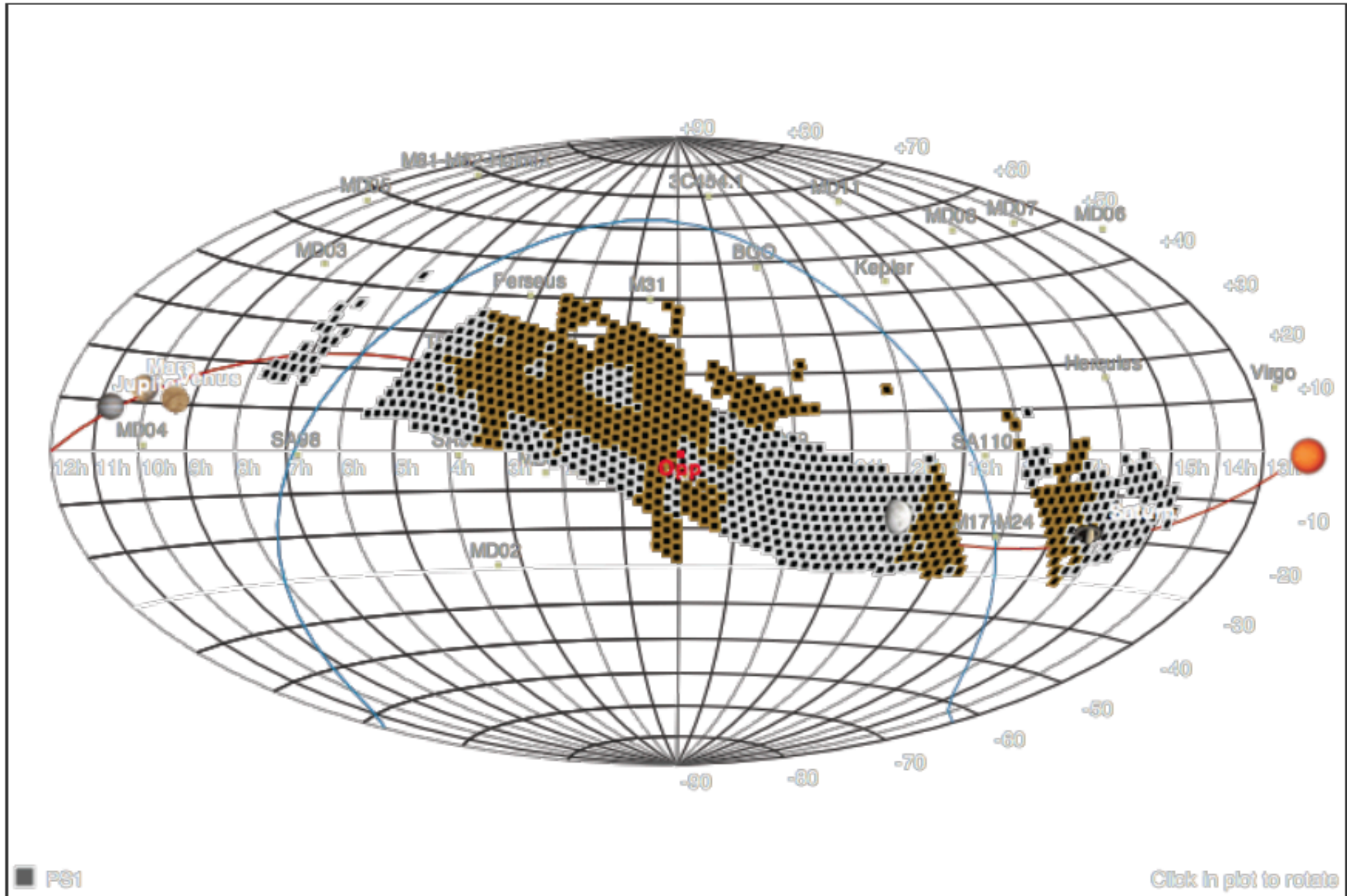
- ✦ The Pan-STARRS1 Science Consortium survey will finish at the end of February 2014
- ✦ We are planning to conduct an NEO optimized survey using 100% of the observing time beginning in March 2014
- ✦ This will enable a much larger area of the sky to be surveyed multiple times per month
- ✦ We can survey further south, reaching to -40°

Digest scores of NEOs observed by PS I

PS1 NEO

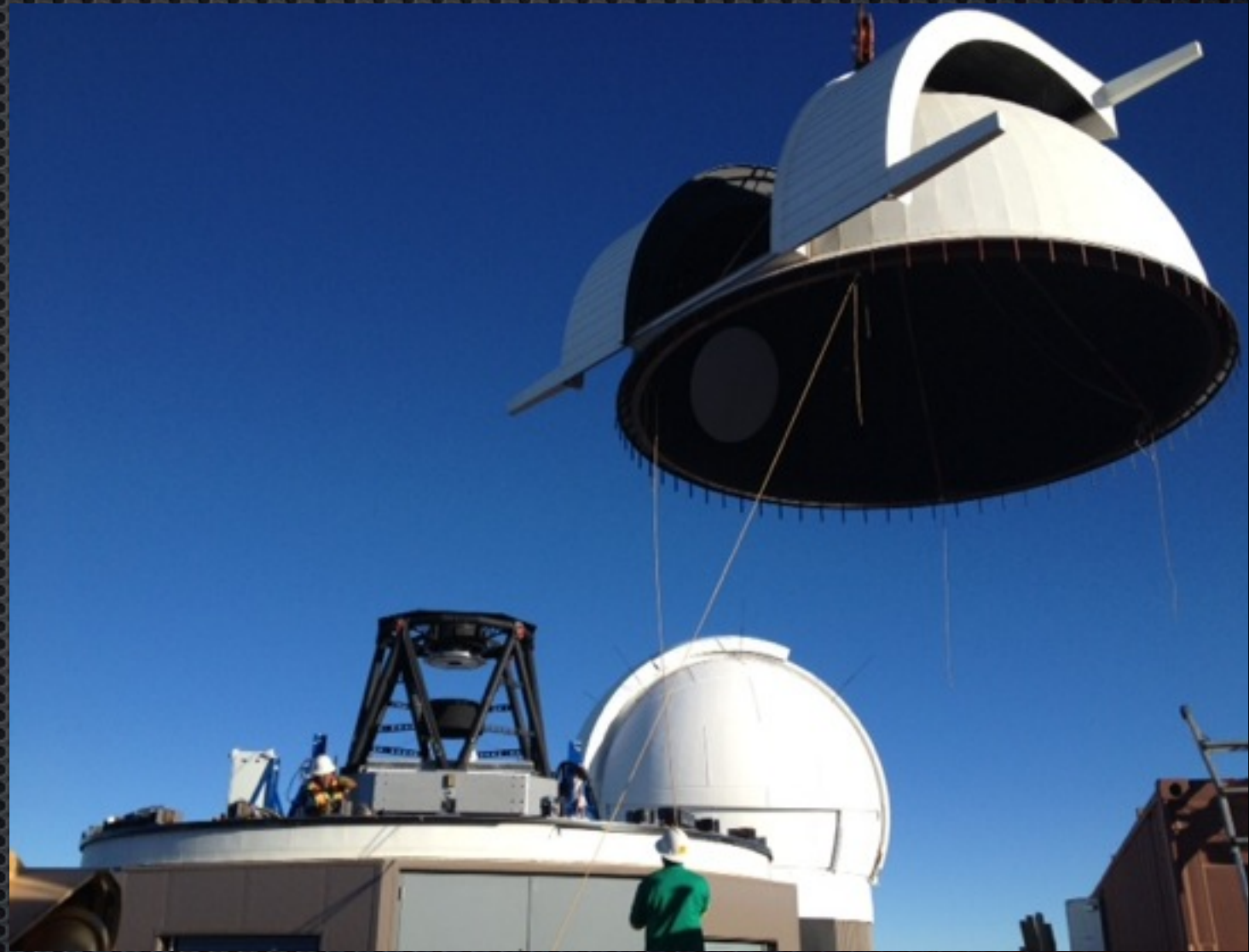


Sky coverage with 100% of Pan-STARRS1



Pan-STARRS2

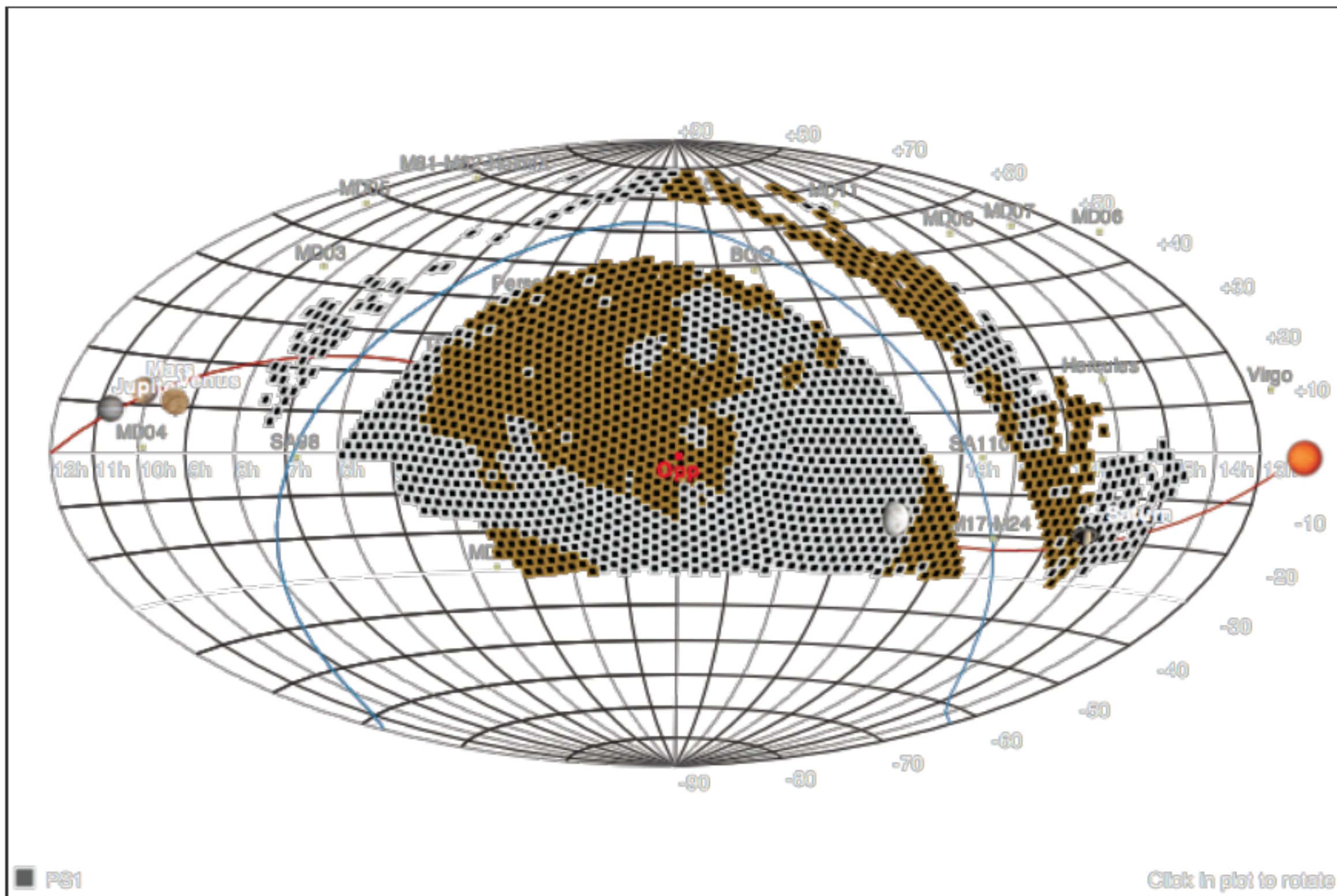
- ✦ Pan-STARRS2 is nearing completion adjacent to PS1
- ✦ Adding a second telescope will enable us to survey a large fraction of the sky multiple times each lunation



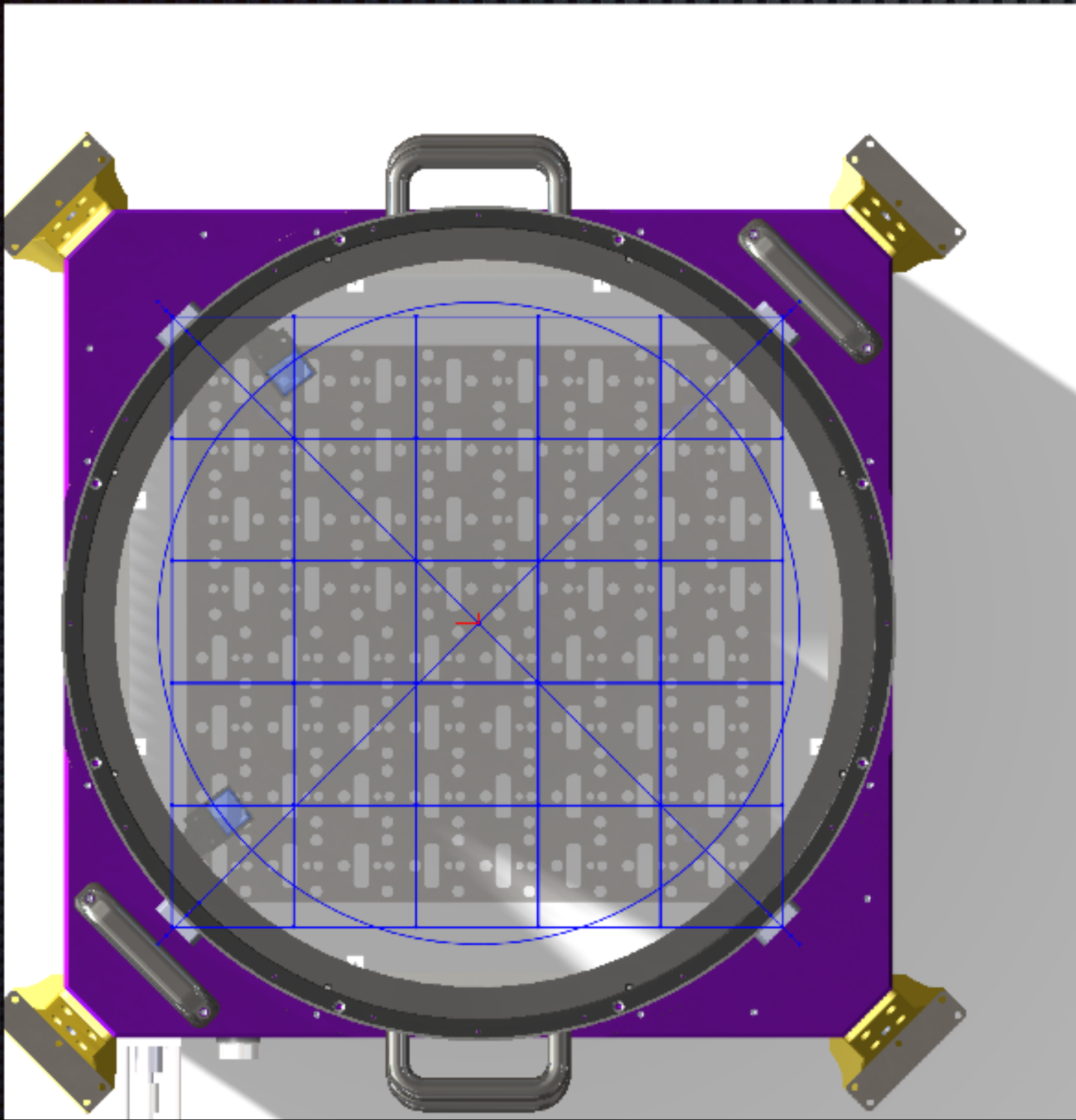
PS2 image of the Crab Nebula



Sky coverage with 100% of Pan-STARRS1&2



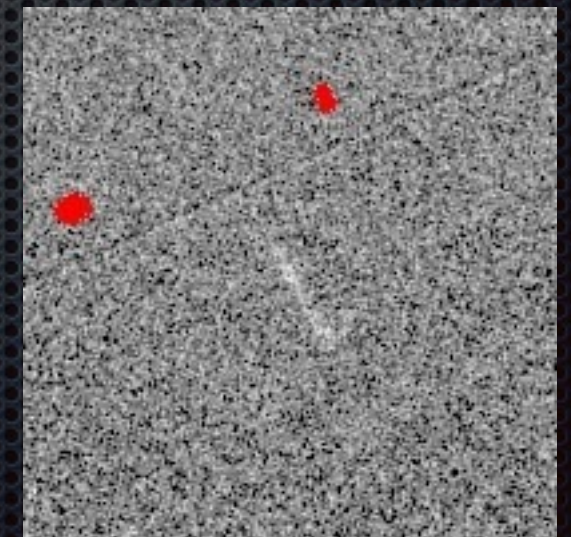
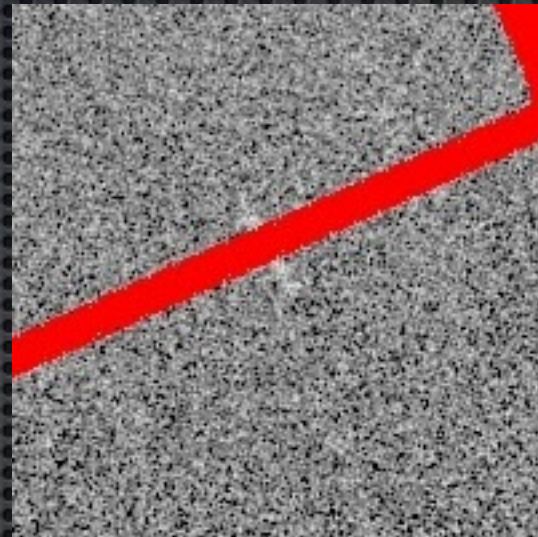
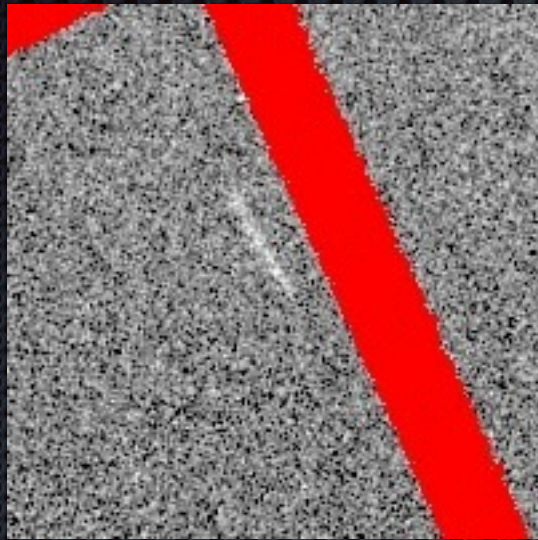
Camera retrofit



The CCDs in the Pan-STARRS telescopes are not ideal

- ✦ Cosmetic quality of some devices is poor
- ✦ Cell structure in CCDs hinders discovery of faster moving NEOs

Example of fast moving asteroid detection compromised by cell gaps



Camera retrofit



- Better CCDs would increase our discovery rate by a factor of almost 2